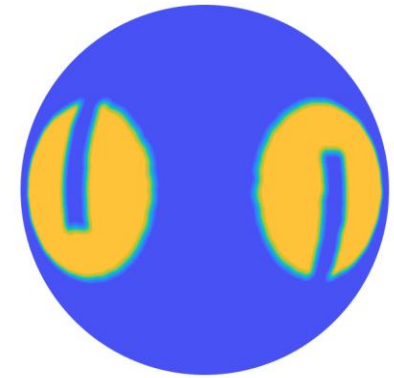
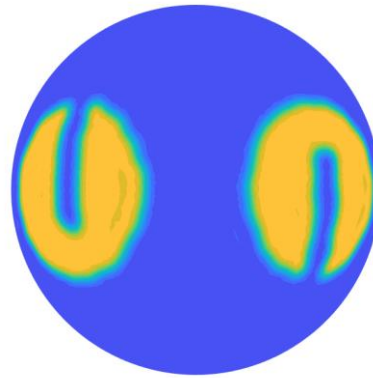
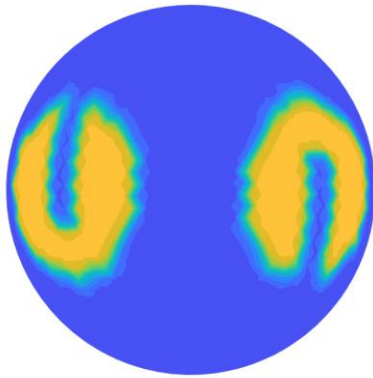


Exceptional service in the national interest



Atmospheric Dynamics with Polyharmonic Spline RBFs

Greg Barnett



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Outline

- Polyharmonic Spline (PHS) RBFs with Polynomials
 - 1D Example
- Interpolation and Differentiation Weights
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 - Weights on the Sphere
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 - Governing Equations
 - Limiter/Fixer
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 - Results
- Eulerian Shallow Water Model
 - Governing Equations
 - Test Cases
 - Results
- Conclusions and Future Work

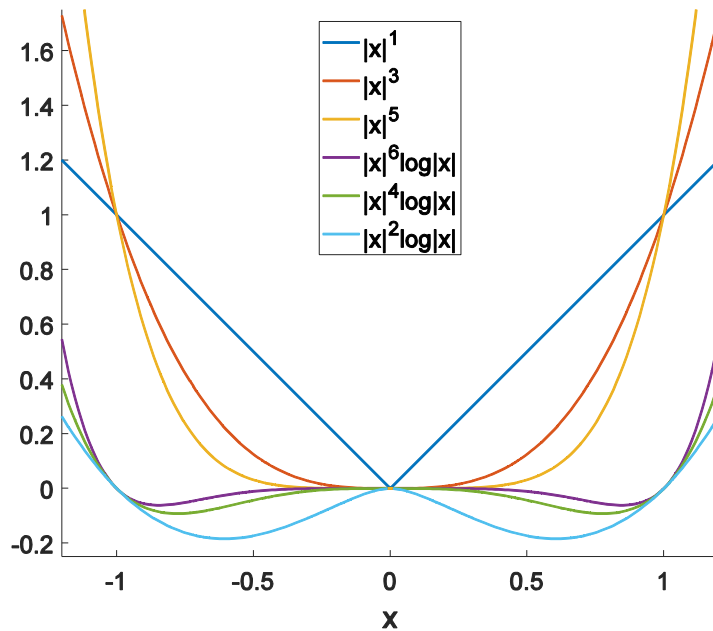
Table of RBFs

$\phi(\mathbf{x}), \mathbf{x} \in \mathbb{R}^d, \varepsilon \in \mathbb{R}$	Name (acronym)
$\sqrt{1 + (\varepsilon \ \mathbf{x}\)^2}$	Multiquadric (MQ)
$\frac{1}{1 + (\varepsilon \ \mathbf{x}\)^2}$	Inverse Quadratic (IQ)
$\frac{1}{\sqrt{1 + (\varepsilon \ \mathbf{x}\)^2}}$	Inverse Multiquadric (IMQ)
$e^{-(\varepsilon \ \mathbf{x}\)^2}$	Gaussian (GA)
$\ \mathbf{x}\ ^{2k+1}$ $\ \mathbf{x}\ ^{2k} \log \ \mathbf{x}\ , k \in \mathbb{N}$	Polyharmonic Spline (PHS)

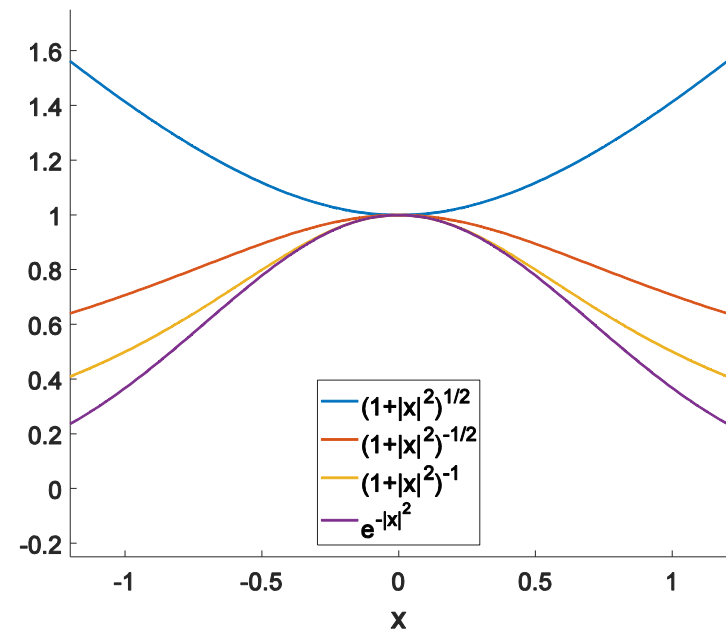
$$\|\cdot\| = \|\cdot\|_2$$

Some RBFs in 1D

Polyharmonic Spline RBFs

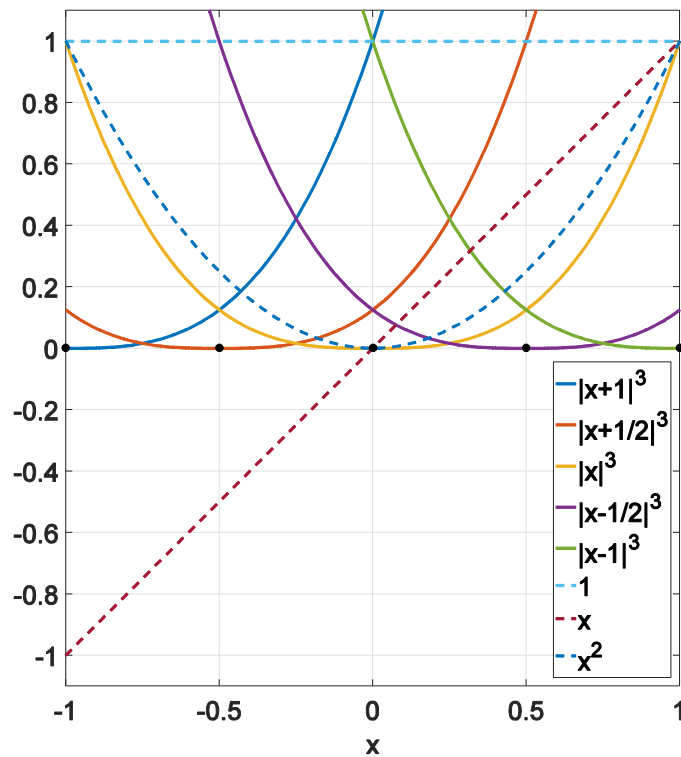


Infinitely Differentiable RBFs

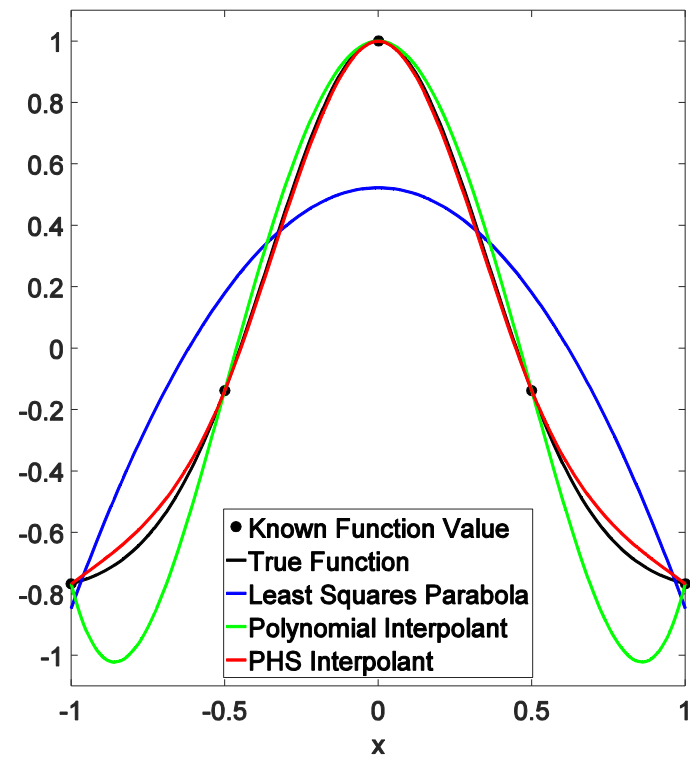


Example: Equi-spaced Interpolation in 1D

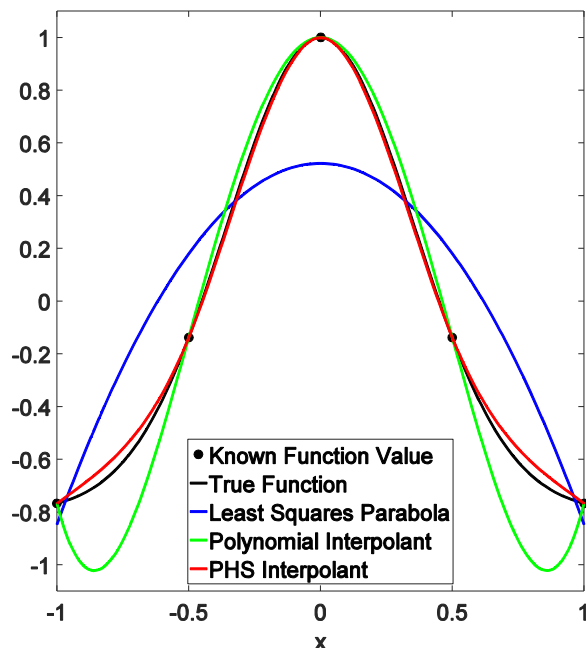
PHS Basis Functions



Approximation



Structure of the Linear System



Least Squares Parabola:

$$\mu_1 + \mu_2 x + \mu_3 x^2$$

$$\begin{bmatrix} 1 & x_1 & x_1^2 \\ 1 & x_2 & x_2^2 \\ 1 & x_3 & x_3^2 \\ 1 & x_4 & x_4^2 \\ 1 & x_5 & x_5^2 \end{bmatrix} \begin{bmatrix} \mu_1 \\ \mu_2 \\ \mu_3 \end{bmatrix} = \begin{bmatrix} f_1 \\ f_2 \\ f_3 \\ f_4 \\ f_5 \end{bmatrix}$$

Polynomial Interpolant:

$$\mu_1 + \mu_2 x + \mu_3 x^2 + \mu_4 x^3 + \mu_5 x^4$$

$$\begin{bmatrix} 1 & x_1 & x_1^2 & x_1^3 & x_1^4 \\ 1 & x_2 & x_2^2 & x_2^3 & x_2^4 \\ 1 & x_3 & x_3^2 & x_3^3 & x_3^4 \\ 1 & x_4 & x_4^2 & x_4^3 & x_4^4 \\ 1 & x_5 & x_5^2 & x_5^3 & x_5^4 \end{bmatrix} \begin{bmatrix} \mu_1 \\ \mu_2 \\ \mu_3 \\ \mu_4 \\ \mu_5 \end{bmatrix} = \begin{bmatrix} f_1 \\ f_2 \\ f_3 \\ f_4 \\ f_5 \end{bmatrix}$$

PHS Interpolant:

$$\sum_{j=1}^5 \lambda_j |x - x_j|^3 + \mu_1 + \mu_2 x + \mu_3 x^2$$

$$\begin{bmatrix} 0 & |x_1 - x_2|^3 & |x_1 - x_3|^3 & |x_1 - x_4|^3 & |x_1 - x_5|^3 \\ |x_2 - x_1|^3 & 0 & |x_2 - x_3|^3 & |x_2 - x_4|^3 & |x_2 - x_5|^3 \\ |x_3 - x_1|^3 & |x_3 - x_2|^3 & 0 & |x_3 - x_4|^3 & |x_3 - x_5|^3 \\ |x_4 - x_1|^3 & |x_4 - x_2|^3 & |x_4 - x_3|^3 & 0 & |x_4 - x_5|^3 \\ |x_5 - x_1|^3 & |x_5 - x_2|^3 & |x_5 - x_3|^3 & |x_5 - x_4|^3 & 0 \\ \hline 1 & 1 & 1 & 1 & 1 \\ x_1 & x_2 & x_3 & x_4 & x_5 \\ x_1^2 & x_2^2 & x_3^2 & x_4^2 & x_5^2 \end{bmatrix} \begin{bmatrix} \lambda_1 \\ \lambda_2 \\ \lambda_3 \\ \lambda_4 \\ \lambda_5 \\ \mu_1 \\ \mu_2 \\ \mu_3 \end{bmatrix} = \begin{bmatrix} f_1 \\ f_2 \\ f_3 \\ f_4 \\ f_5 \\ 0 \\ 0 \\ 0 \end{bmatrix}$$

Properties of Polyharmonic Splines

- PHS basis includes both RBFs and polynomials
 - RBFs improve performance and allow the use of irregular nodes
 - polynomials give convergence to smooth solutions (no saturation error)
- The interpolation problem is guaranteed to have a unique solution provided that polynomials are included up to the required degree, and the nodes are unisolvent. For $k = 1, 2, 3, \dots$
 - $\phi(x) = \|x\|^{2k} \log \|x\|$
 - Polynomials up to degree k or higher
 - $\phi(x) = \|x\|^{2k+1}$
 - Polynomials up to degree k or higher
 - Rule of thumb for modest polynomial degrees: Twice as many RBFs as polynomials
- Condition number of PHS A -matrix is invariant under rotation, translation, and uniform scaling
- No need to search for optimal shape parameter

Interpolation in 2D [$\mathbf{x} = (x, y)$]

Given nodes $\{(x_i, y_i)\}_{i=1}^n$ and corresponding function values $\{f_i\}_{i=1}^n$, find a linear combination of RBF and polynomial basis functions that matches the data exactly.

1. Assume the appropriate form of the underlying approximation:

$$\Phi(x, y) = \sum_{j=1}^n \lambda_j \phi_j(x, y) + \sum_{k=1}^m \mu_k p_k(x, y),$$

where $\phi_j(x, y) = \phi(x - x_j, y - y_j)$.

2. Require Φ to match the data at each node:

$$\Phi(x_i, y_i) = \sum_{j=1}^n \lambda_j \phi_j(x_i, y_i) + \sum_{k=1}^m \mu_k p_k(x_i, y_i) = f_i, \quad i = 1, 2, 3, \dots, n.$$

3. Enforce regularity conditions on the coefficients $\{\lambda_j\}$:

$$\sum_{j=1}^n \lambda_j p_k(x_j, y_j) = 0, \quad k = 1, 2, 3, \dots, m.$$

4. Solve the symmetric linear system for $\{\lambda_j\}$ and $\{\mu_k\}$.

Differentiation Weights in 2D

Interpolation Problem:

$$\begin{bmatrix} \mathbf{A} & \mathbf{P} \\ \mathbf{P}^T & \mathbf{O} \end{bmatrix} \begin{bmatrix} \boldsymbol{\lambda} \\ \boldsymbol{\mu} \end{bmatrix} = \begin{bmatrix} \mathbf{f} \\ \mathbf{O} \end{bmatrix},$$

$$\begin{aligned} a_{ij} &= \phi_j(x_i, y_i) = \phi(x_i - x_j, y_i - y_j), & i, j &= 1, 2, 3, \dots, n, \\ p_{ik} &= p_k(x_i, y_i), & i &= 1, 2, 3, \dots, n, & k &= 1, 2, 3, \dots, m. \end{aligned}$$

Use $[L\Phi](\tilde{x}, \tilde{y})$ to approximate $[Lf](\tilde{x}, \tilde{y})$:

$$\begin{aligned} [L\Phi](\tilde{x}, \tilde{y}) &= \sum_{j=1}^n \lambda_j [L\phi_j](\tilde{x}, \tilde{y}) + \sum_{k=1}^m \mu_k [Lp_k](\tilde{x}, \tilde{y}) \\ &= [\mathbf{b} \quad \mathbf{c}] \begin{bmatrix} \boldsymbol{\lambda} \\ \boldsymbol{\mu} \end{bmatrix} = \underbrace{\left([\mathbf{b} \quad \mathbf{c}] \begin{bmatrix} \mathbf{A} & \mathbf{P} \\ \mathbf{P}^T & \mathbf{O} \end{bmatrix}^{-1} \right)}_{\text{weights}} \begin{bmatrix} \mathbf{f} \\ \mathbf{O} \end{bmatrix}, \end{aligned}$$

$$\begin{aligned} b_j &= [L\phi_j](\tilde{x}, \tilde{y}), & j &= 1, 2, 3, \dots, n, \\ c_k &= [Lp_k](\tilde{x}, \tilde{y}), & k &= 1, 2, 3, \dots, m. \end{aligned}$$

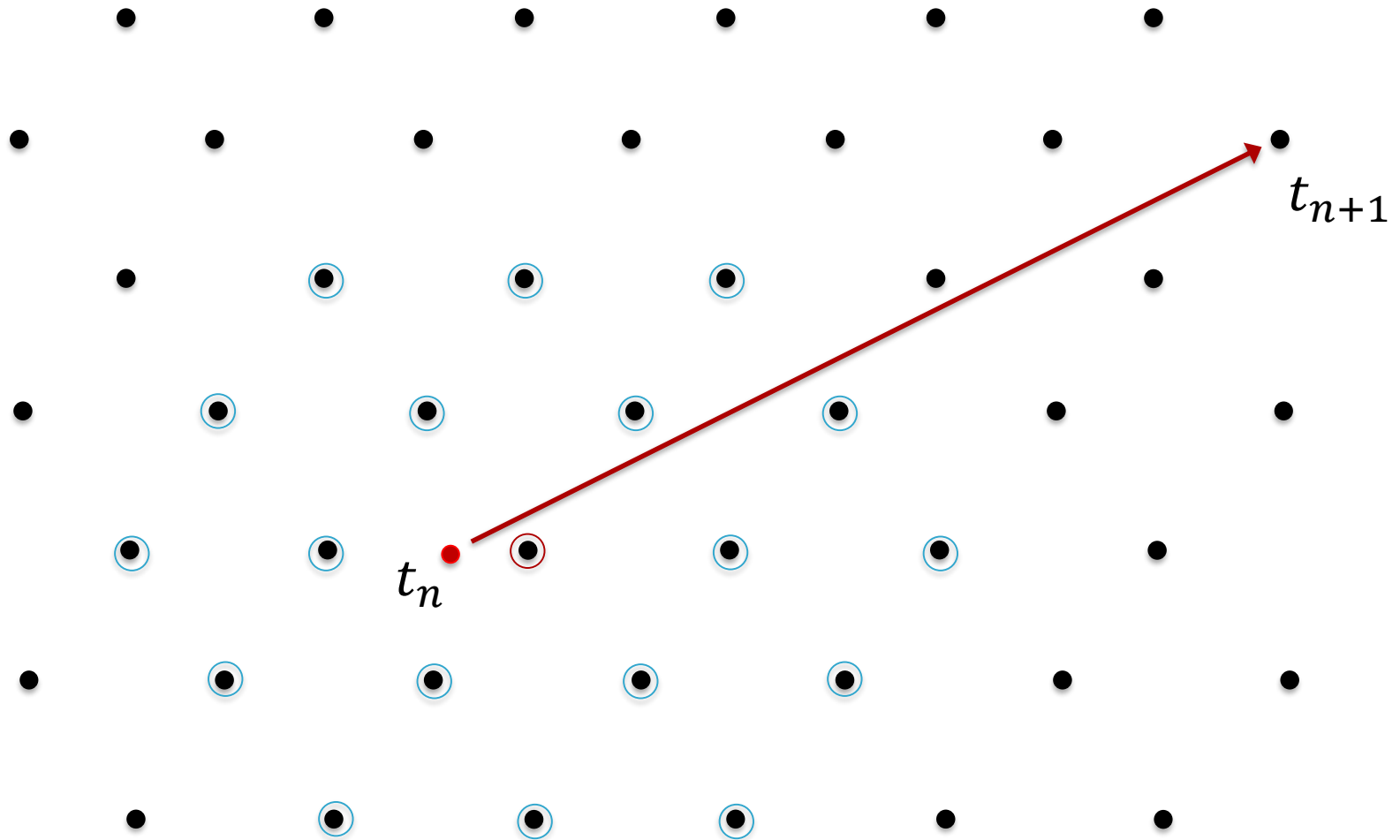
Derivative Approximation on the Sphere

- Given: nodes $[\mathbf{x}, \mathbf{y}, \mathbf{z}]$ on the sphere and function values f
- Find: differentiation matrices (DMs) $\mathbf{W}_x, \mathbf{W}_y, \mathbf{W}_z$ to approximate $\frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z}$
- Method: Use the fact that $\nabla f(\tilde{x}, \tilde{y}, \tilde{z})$ is tangent to the sphere at $(\tilde{x}, \tilde{y}, \tilde{z})$
 - For each node, get orthogonal unit vectors $\mathbf{e}_{\hat{x}}$ and $\mathbf{e}_{\hat{y}}$ tangent to the sphere
 - Use 2D method to get matrices $\mathbf{W}_{\hat{x}}$ and $\mathbf{W}_{\hat{y}}$ that approximate $\frac{\partial}{\partial \hat{x}}$ and $\frac{\partial}{\partial \hat{y}}$
 - $\nabla f = \mathbf{e}_{\hat{x}} \frac{\partial f}{\partial \hat{x}} + \mathbf{e}_{\hat{y}} \frac{\partial f}{\partial \hat{y}}$
 - $\frac{\partial f}{\partial x} = (\nabla f)_1 = (\mathbf{e}_{\hat{x}})_1 \frac{\partial f}{\partial \hat{x}} + (\mathbf{e}_{\hat{y}})_1 \frac{\partial f}{\partial \hat{y}} = \left\{ (\mathbf{e}_{\hat{x}})_1 \frac{\partial}{\partial \hat{x}} + (\mathbf{e}_{\hat{y}})_1 \frac{\partial}{\partial \hat{y}} \right\} f$
 - $\mathbf{W}_x = \text{diag}\{(\mathbf{e}_{\hat{x}})_1\} \mathbf{W}_{\hat{x}} + \text{diag}\{(\mathbf{e}_{\hat{y}})_1\} \mathbf{W}_{\hat{y}}$
 - $\mathbf{W}_y = \text{diag}\{(\mathbf{e}_{\hat{x}})_2\} \mathbf{W}_{\hat{x}} + \text{diag}\{(\mathbf{e}_{\hat{y}})_2\} \mathbf{W}_{\hat{y}}$
 - $\mathbf{W}_z = \text{diag}\{(\mathbf{e}_{\hat{x}})_3\} \mathbf{W}_{\hat{x}} + \text{diag}\{(\mathbf{e}_{\hat{y}})_3\} \mathbf{W}_{\hat{y}}$

Semi-Lagrangian Transport

- Governing Equations (velocity \mathbf{u} is a known function)
 - $\frac{\partial \rho}{\partial t} = -\mathbf{u} \cdot \nabla \rho - \rho \nabla \cdot \mathbf{u},$ (Eulerian, short time-steps)
 - $\frac{Dq}{Dt} = \left(\frac{\partial}{\partial t} + \mathbf{u} \cdot \nabla \right) q = 0.$ (Semi-Lagrangian, long time-steps)
- Quasi-Monotone Limiter for q (const. along flow trajectories)
 - $m_k = \min_{\ell} \{q_{k_{\ell}}^{(n)}\}$ and $M_k = \max_{\ell} \{q_{k_{\ell}}^{(n)}\}$
 - Set $q_k^{(n+1)} = \min \{q_k^{(n+1)}, M_k\}$
 - Set $q_k^{(n+1)} = \max \{q_k^{(n+1)}, m_k\}$
- Mass Fixer ($\text{tracerMass} \equiv \sum_{k=1}^N \rho_k q_k V_k$)
 - If $\text{tracerMass} < \text{initialMass}$, add mass in cells with $q_k < M_k$
 - If $\text{tracerMass} > \text{initialMass}$, subtract mass from cells with $q_k > m_k$

Semi-Lagrangian Transport



Pre-Processing

- Get DMs \mathbf{W}_x , \mathbf{W}_y , and \mathbf{W}_z for the continuity equation
- Find index of the n nearest neighbors to each node
- For each node $\mathbf{x}_k = (x_k, y_k, z_k)$, write its n neighbors in terms of two orthogonal unit vectors $(\mathbf{e}_{\hat{x}}, \mathbf{e}_{\hat{y}})$ tangent to the sphere at \mathbf{x}_k , and one unit vector $\mathbf{e}_{\hat{z}}$ normal to the sphere at \mathbf{x}_k , so that

$$\mathbf{x}_{k_\ell} = \hat{x}_{k_\ell} \mathbf{e}_{\hat{x}} + \hat{y}_{k_\ell} \mathbf{e}_{\hat{y}} + \hat{z}_{k_\ell} \mathbf{e}_{\hat{z}}, \quad \ell = 1, 2, 3, \dots, n.$$

- Set $\mathbf{C}_k = \begin{bmatrix} \mathbf{A}_k & \mathbf{P}_k \\ \mathbf{P}_k^T & \mathbf{O}_{6 \times 6} \end{bmatrix}^{-1} \begin{bmatrix} \mathbf{I}_n \\ \mathbf{O}_{6 \times n} \end{bmatrix}$, where

$$(\mathbf{A}_k)_{ij} = \phi \left(\hat{x}_{k_i} - \hat{x}_{k_j}, \hat{y}_{k_i} - \hat{y}_{k_j} \right), \quad i, j = 1, 2, 3, \dots, n,$$

$$\mathbf{P}_k = [\mathbf{1}, \hat{\mathbf{x}}_k, \hat{\mathbf{y}}_k, \hat{\mathbf{x}}_k^2, \hat{\mathbf{x}}_k \hat{\mathbf{y}}_k, \hat{\mathbf{y}}_k^2].$$

Time Stepping

1. Step $\frac{\partial \rho}{\partial t} = -\mathbf{u} \cdot \nabla \rho - \rho \nabla \cdot \mathbf{u}$ from t_n to t_{n+1} using several explicit Eulerian time steps (RK3)
2. Step $\mathbf{x}' = -\mathbf{u}$ from t_{n+1} to t_n to get departure points (RK4)
3. Find the nearest fixed neighbor to each departure point
4. Use the corresponding pre-calculated cardinal coefficients $\{\mathbf{C}_k\}$ and the newly formed row-vectors $\{\mathbf{b}_k\}$ to get rows of the interpolation matrix \mathbf{W} ($\mathbf{W}_{k\cdot} = \mathbf{b}_k \mathbf{C}_k$)
5. Update q on fixed nodes using weights \mathbf{W}
6. Cycle quasi-monotone limiter and mass-fixer until the tracer mass is nearly equal to the initial tracer mass (diff<1e-13)
7. Repeat

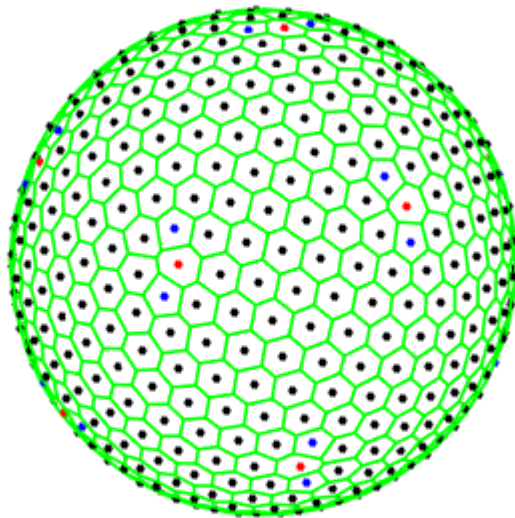
Hyperviscosity

- Add a small dissipative term to the continuity equation
 - $\frac{\partial \rho}{\partial t} = -\mathbf{u} \cdot \nabla \rho - \rho \nabla \cdot \mathbf{u} + \gamma \max\|\mathbf{u}\| (\Delta x)^{2K-1} \Delta^K \rho$
- Reduce high-frequency noise while keeping order of convergence intact
- Achieve stability in time using explicit time-stepping
- PHS are ideal for hyperviscosity, because applying the Laplace operator to a PHS returns another PHS
 - $\phi(x, y) = (x^2 + y^2)^{m/2}$
 - $[\Delta \phi](x, y) = \frac{\partial^2 \phi}{\partial x^2} + \frac{\partial^2 \phi}{\partial y^2} = m^2 (x^2 + y^2)^{(m-2)/2}$
- Parameter $\gamma \in \mathbb{R}$ is determined experimentally at low resolution, and remains unchanged as resolution increases

Transport Test Cases on the Sphere

- Initial Condition q_0 ($\rho_0 = 1$ in all cases)
 - Taken from Nair and Lauritzen, 2010 (NL2010)
 - Gaussian Hills (infinitely differentiable)
 - Cosine Bells (once continuously differentiable)
 - Slotted Cylinders (not continuous)
- Velocity Field (NL2010)
 - Case 1: Translating, vorticity-dominated flow ($\text{CFL}_{\max} = \frac{\max\|\mathbf{u}\|\Delta t}{\Delta x} \approx 8$)
 - Case 2: Translating, divergence-dominated flow ($\text{CFL}_{\max} \approx 5$)
- Spatial Approximations (Minimum Energy (ME) Nodes)
 - Number of nodes $N = 24^2(576), 48^2(2304), 96^2(9216), 192^2(36864)$
 - Interpolation (semi-Lagrangian tracer transport)
 - $\|\mathbf{x}\|^3 + p1 + n19$
 - Derivative Approximation (Eulerian continuity equation)
 - $\|\mathbf{x}\|^2 \log\|\mathbf{x}\| + p5 + n42$
- Time-stepping from $t = 0$ to $t = 5$ (one revolution)

Nodes and Initial Conditions for q

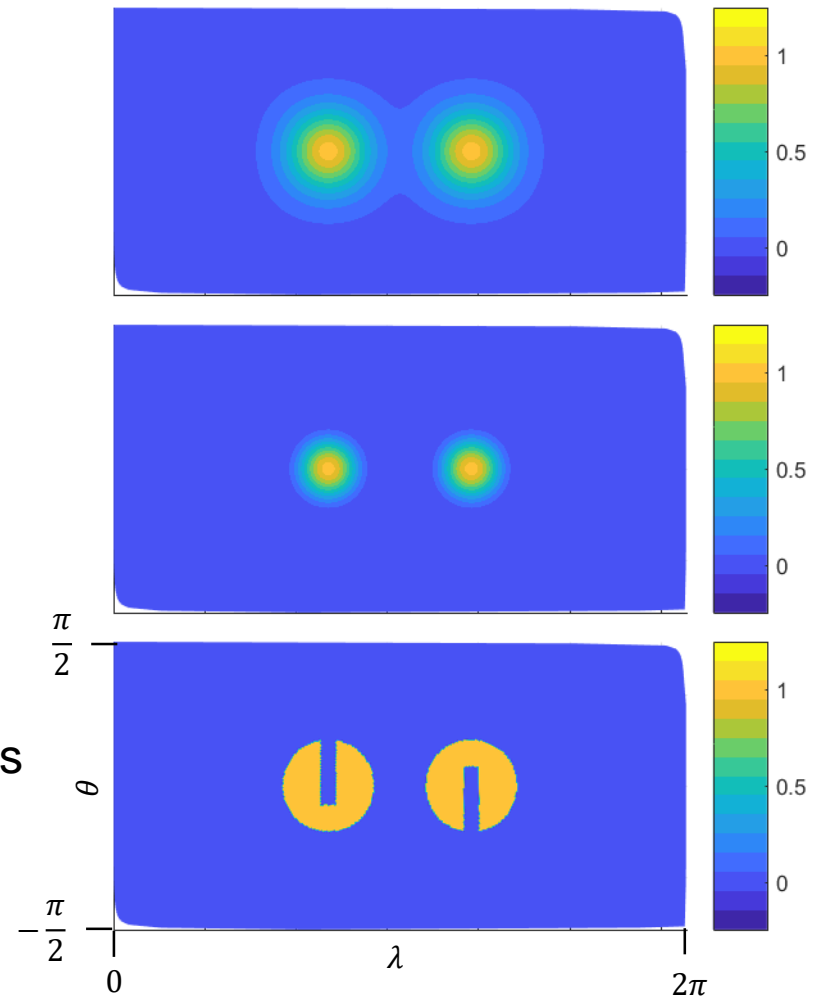


$N = 24^2 = 576$
ME nodes

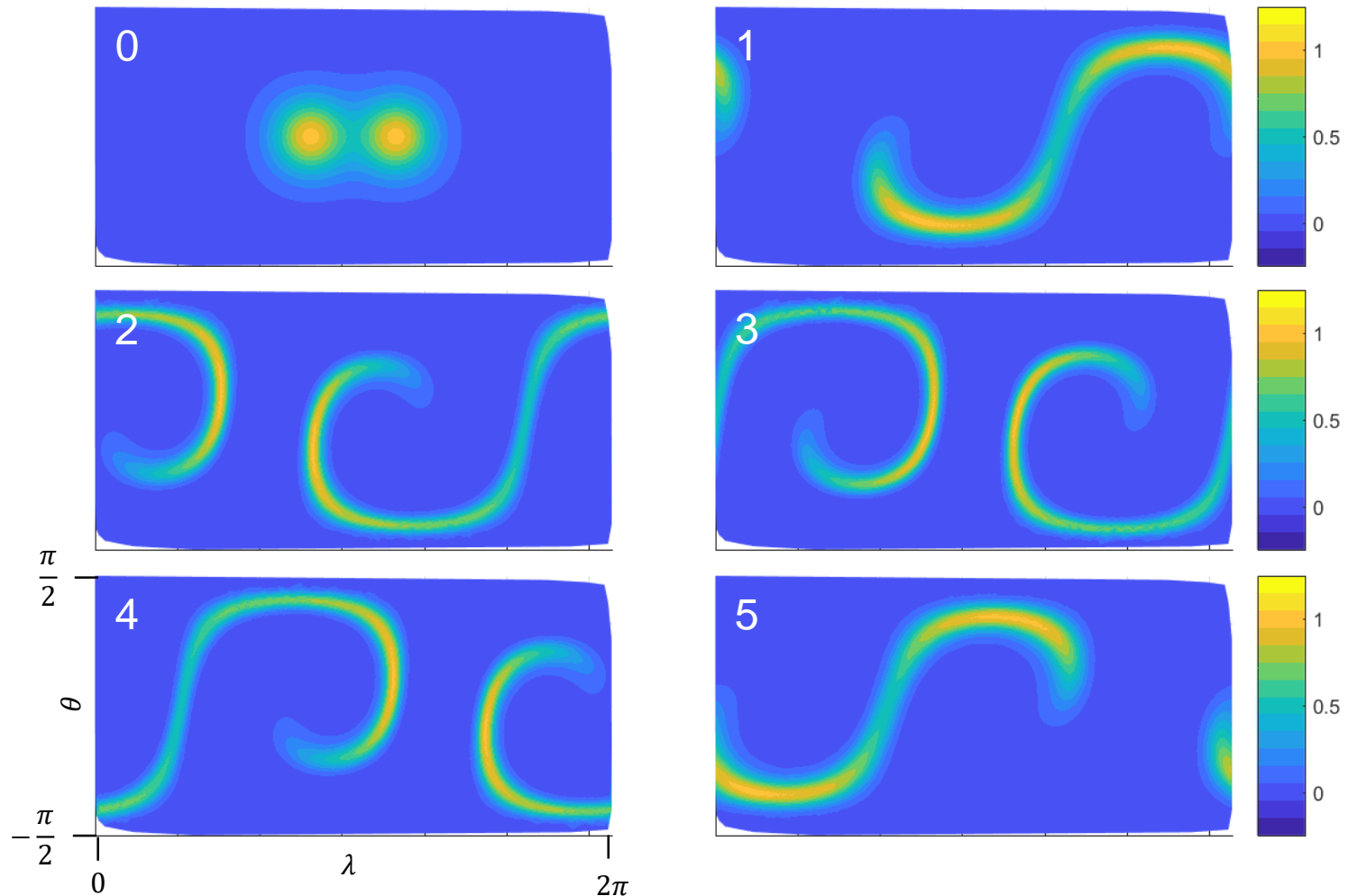
Gaussian Hills
(GH)

Cosine Bells
(CB)

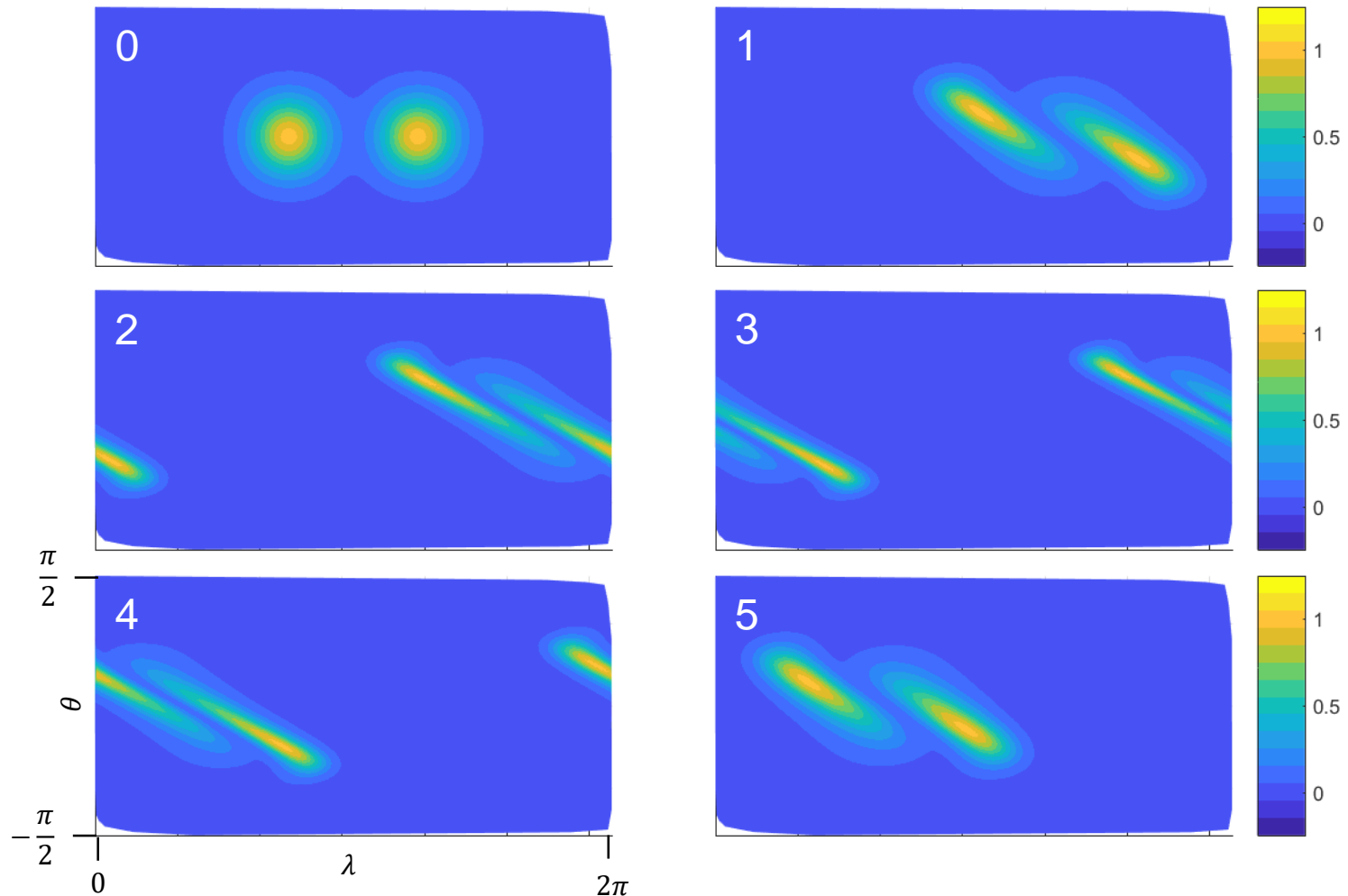
Slotted Cylinders
(SC)



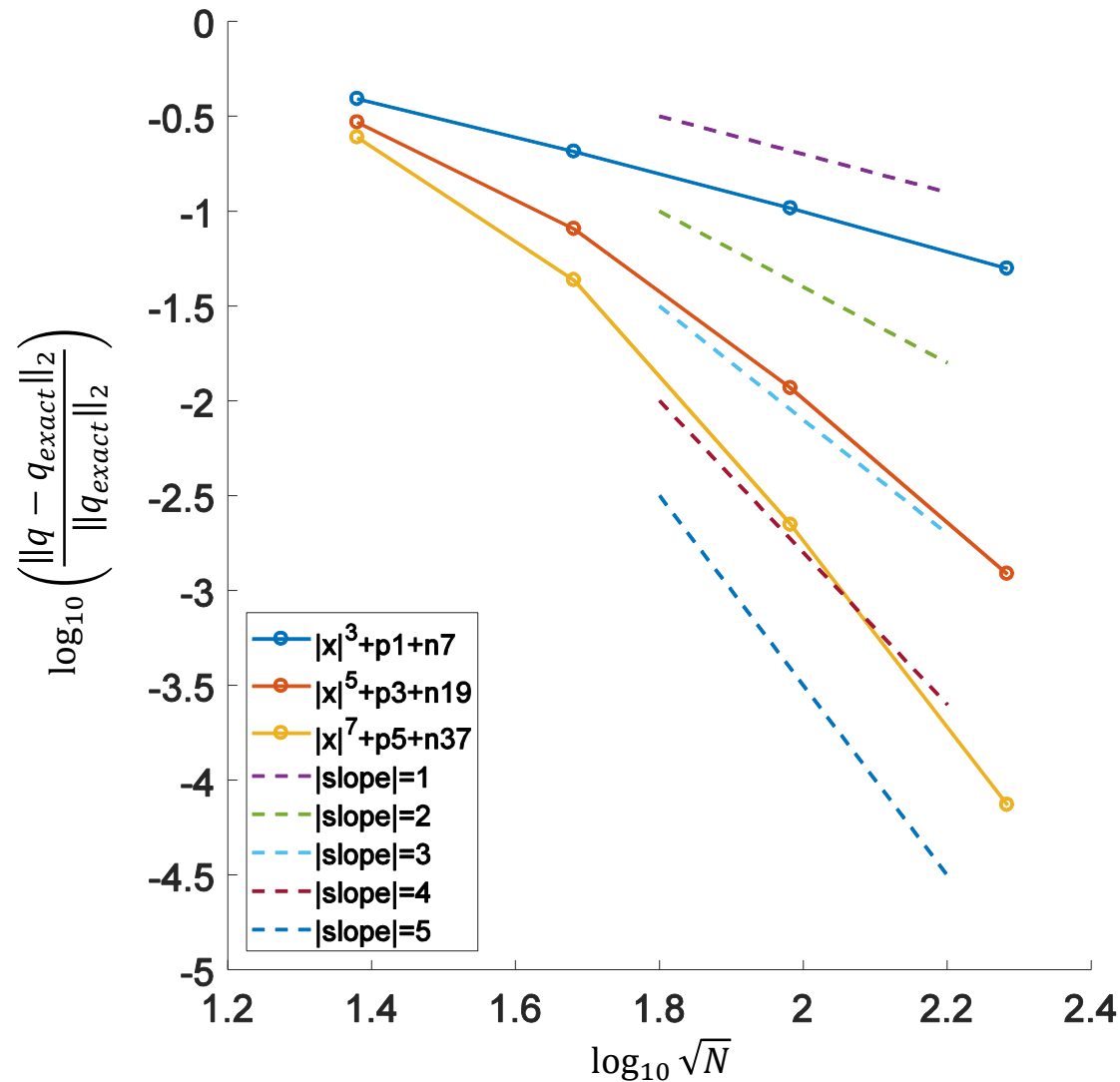
Time Snapshots, Velocity Case 1



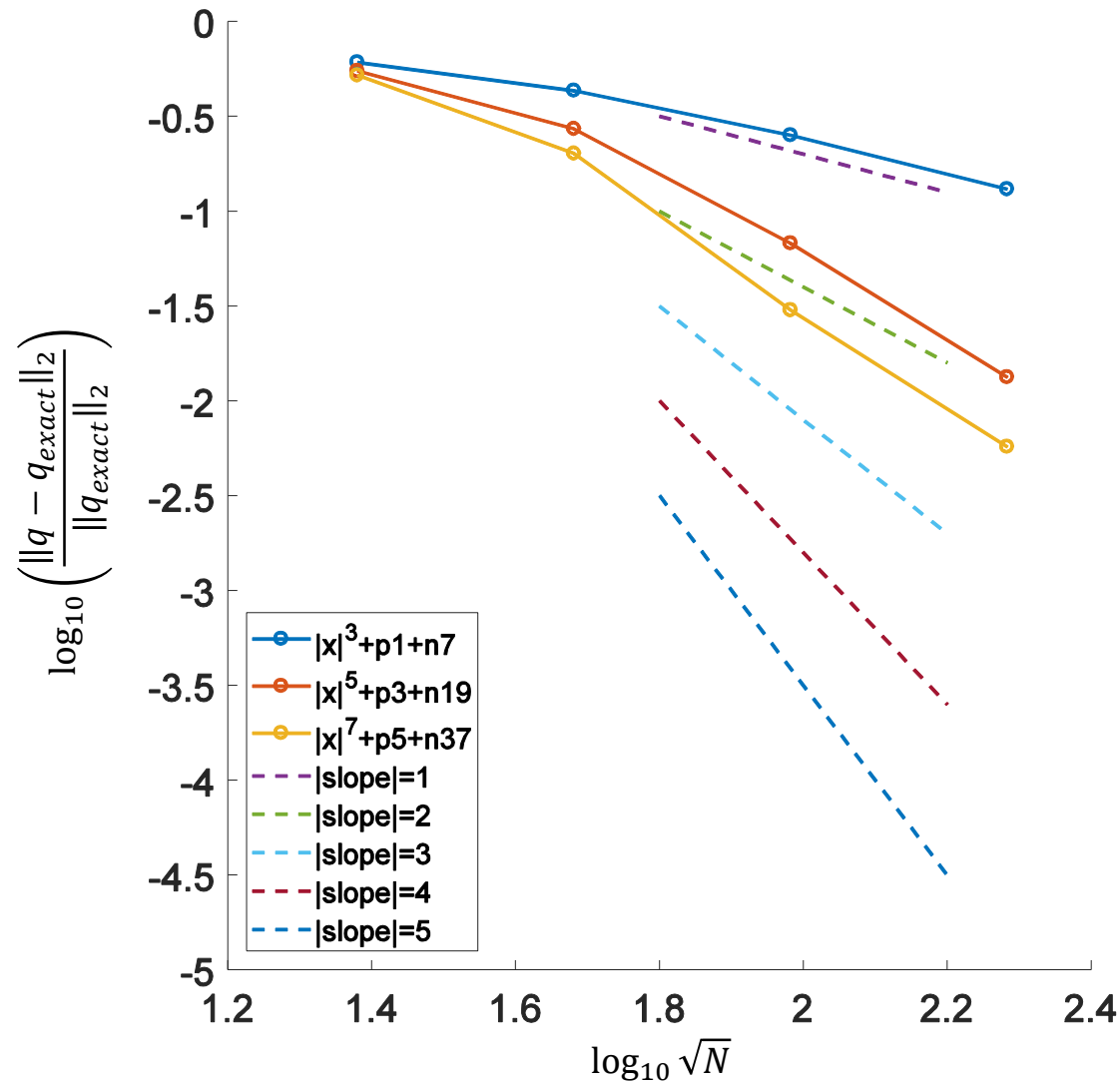
Time Snapshots, Velocity Case 2



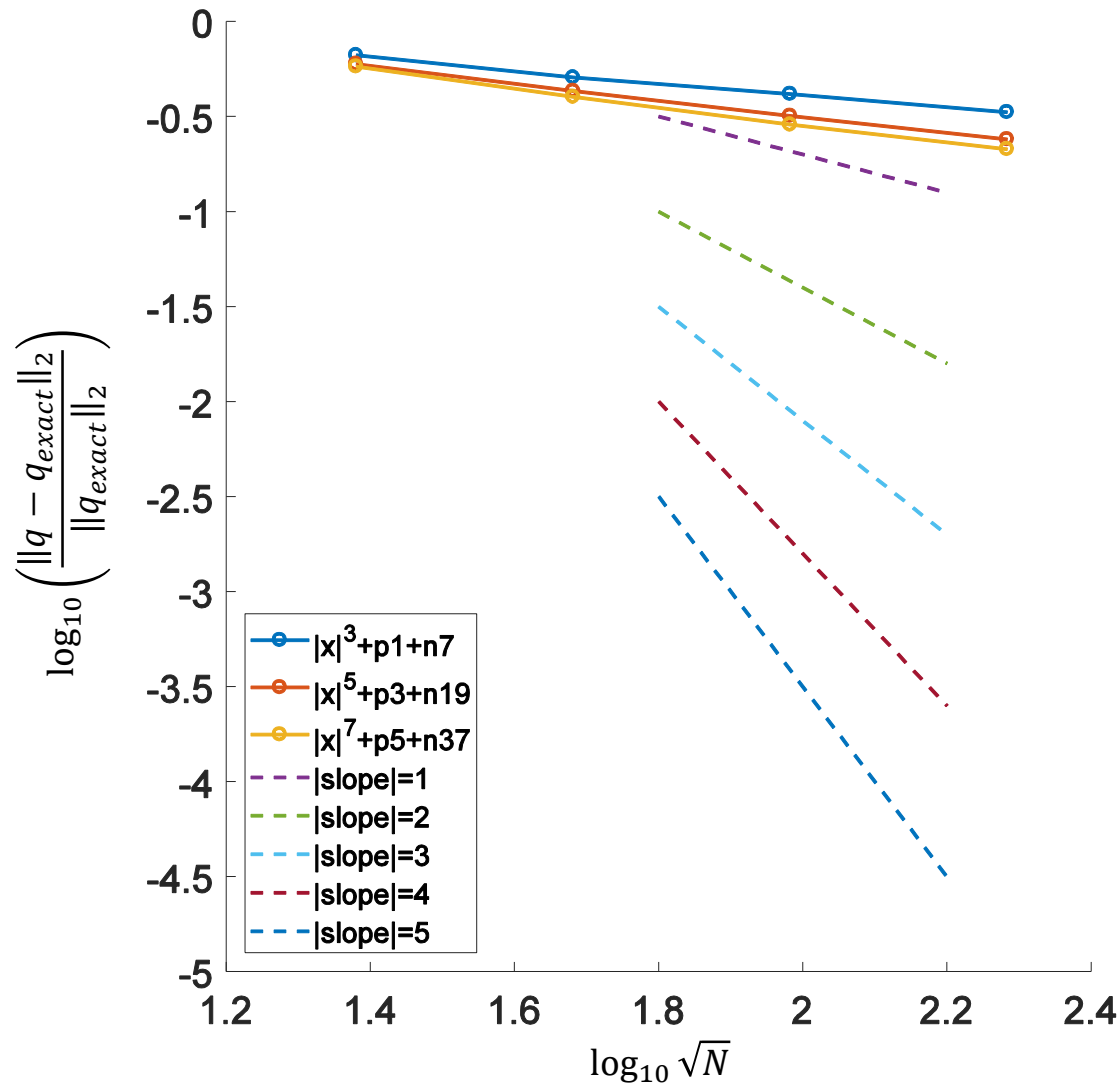
GH, Velocity Case 1, Unlimited



CB, Velocity Case 1, Unlimited



SC, Velocity Case 1, Unlimited



GH, Velocity Case 1

Unlimited

Limited

$N = 48^2$
 $\approx 4.4^\circ$

min=-0.03600780501300, max=0.93671187478677

min=0.00002984696951, max=0.92429725643002

$N = 96^2$
 $\approx 2.2^\circ$

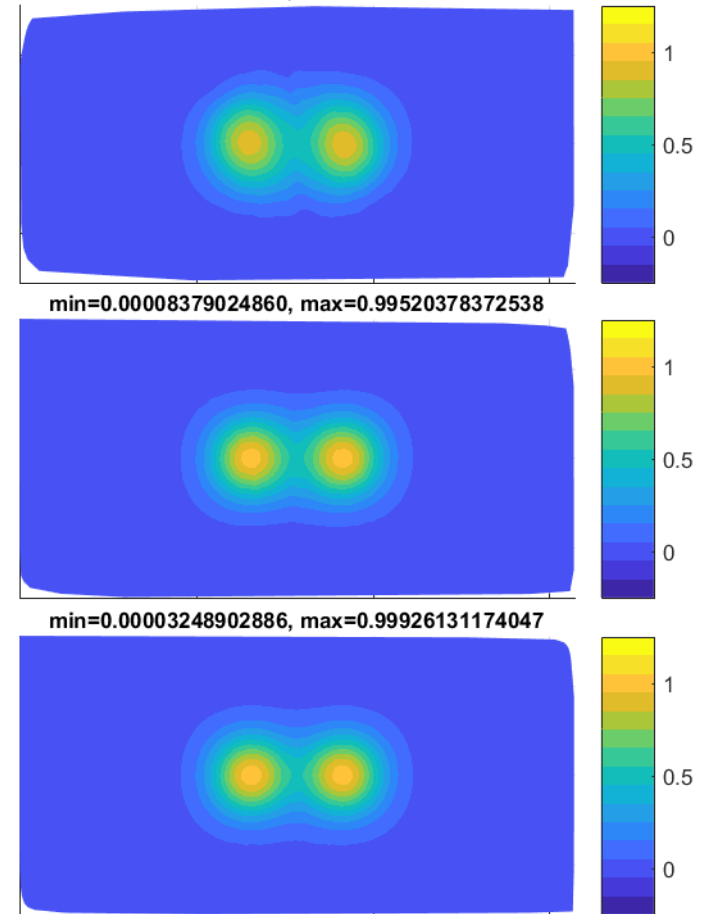
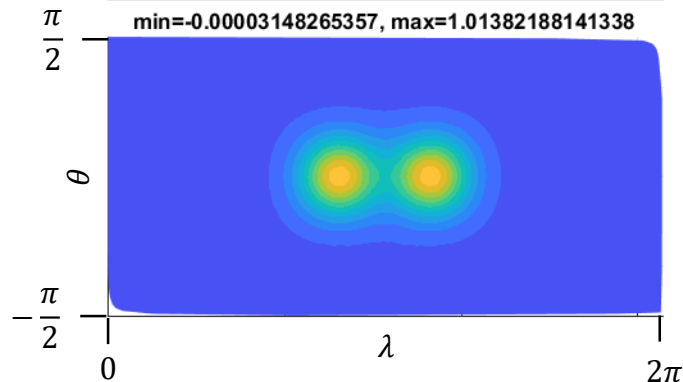
min=-0.00249482626257, max=1.01588345127470

min=0.00008379024860, max=0.99520378372538

$N = 192^2$
 $\approx 1.1^\circ$

min=-0.00003148265357, max=1.01382188141338

min=0.00003248902886, max=0.99926131174047



CB, Velocity Case 1

Unlimited

Limited

$N = 48^2$
 $\approx 4.4^\circ$

min=-0.08739536942368, max=0.78921218944456

min=0.00000000000000, max=0.72100213349540

$N = 96^2$
 $\approx 2.2^\circ$

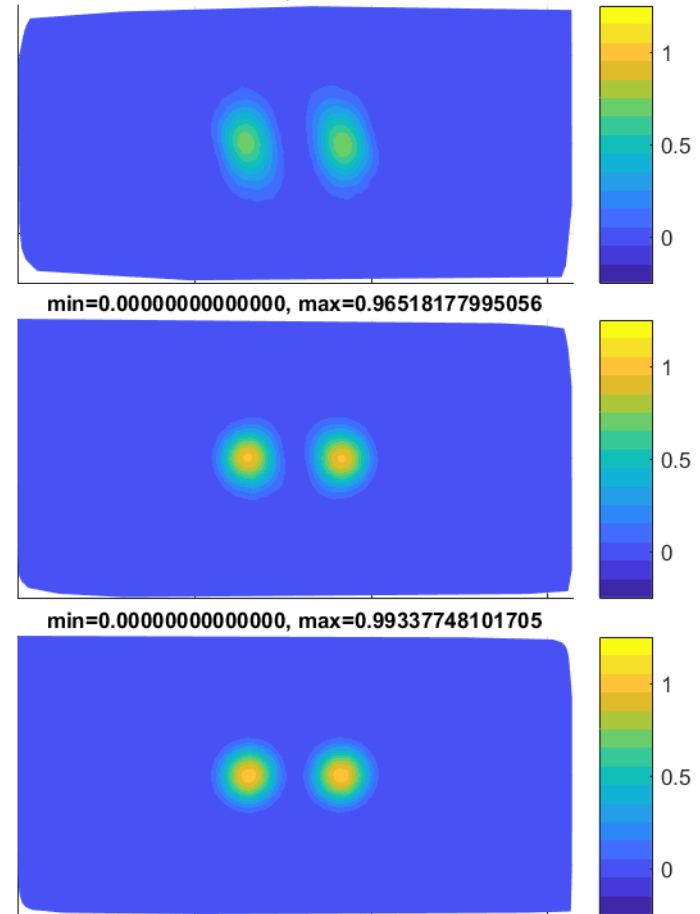
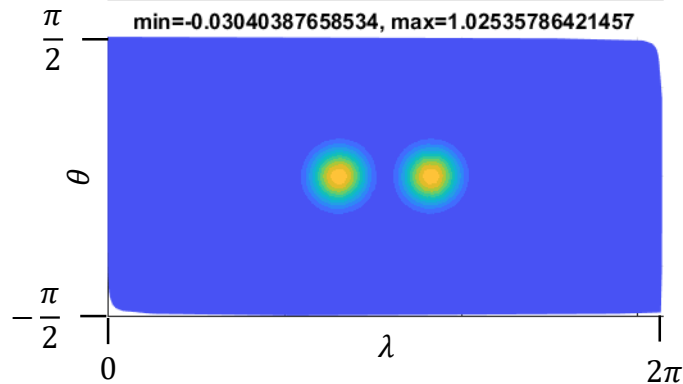
min=-0.05340099181696, max=1.01506032186911

min=0.00000000000000, max=0.96518177995056

$N = 192^2$
 $\approx 1.1^\circ$

min=-0.03040387658534, max=1.02535786421457

min=0.00000000000000, max=0.99337748101705



SC, Velocity Case 1

Unlimited

Limited

$N = 48^2$
 $\approx 4.4^\circ$

min=-0.13362725179515, max=1.19901924844517

min=0.00000000000000, max=0.96860295202275

$N = 96^2$
 $\approx 2.2^\circ$

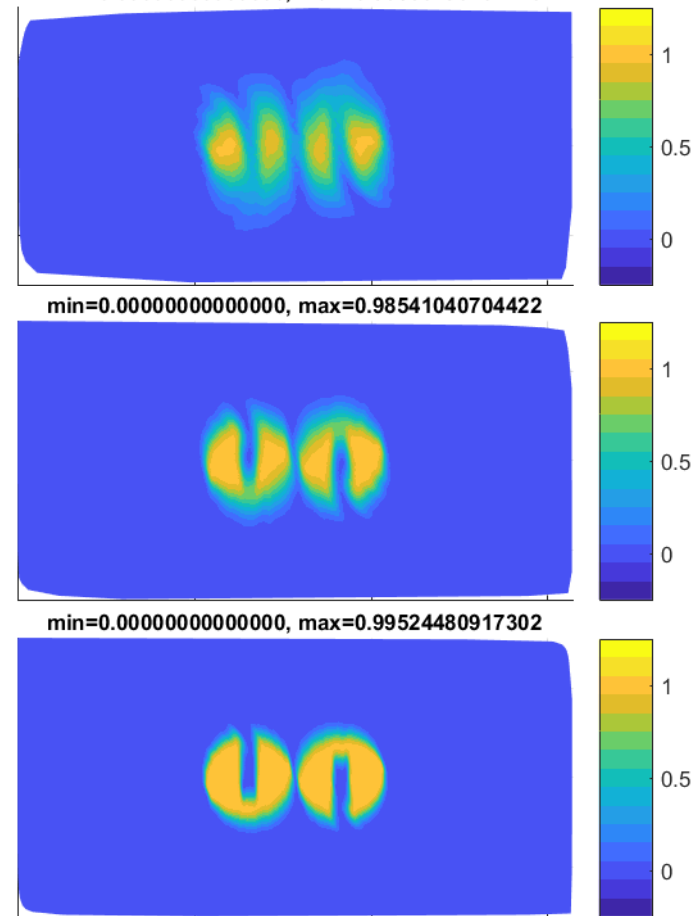
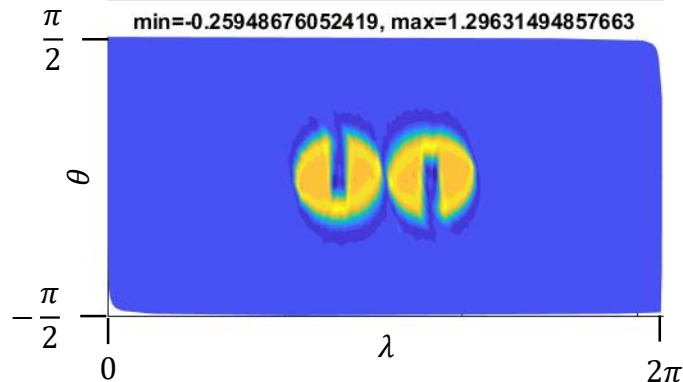
min=-0.20670256411399, max=1.24840235689522

min=0.00000000000000, max=0.98541040704422

$N = 192^2$
 $\approx 1.1^\circ$

min=-0.25948676052419, max=1.29631494857663

min=0.00000000000000, max=0.99524480917302



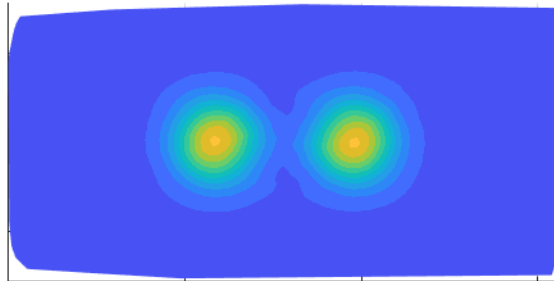
GH, Velocity Case 2

Unlimited

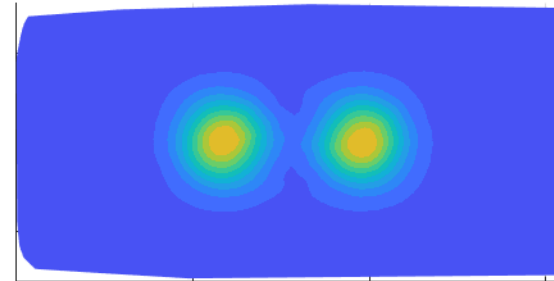
Limited

$N = 48^2$
 $\approx 4.4^\circ$

min=0.00000007808493, max=0.96999177983855

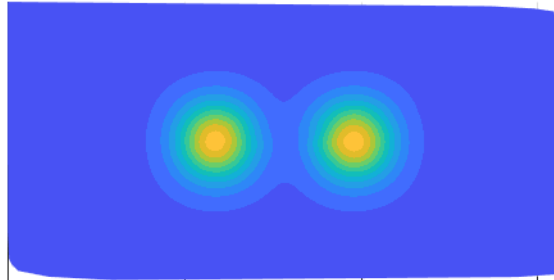


min=0.00001197427628, max=0.94904534685949

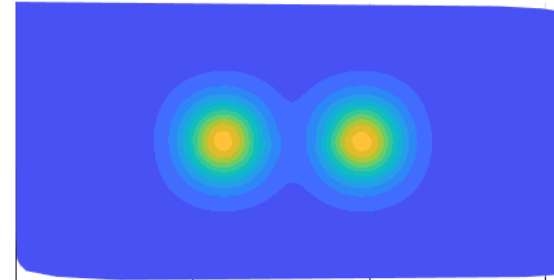


$N = 96^2$
 $\approx 2.2^\circ$

min=0.00000007571221, max=1.01508973972181

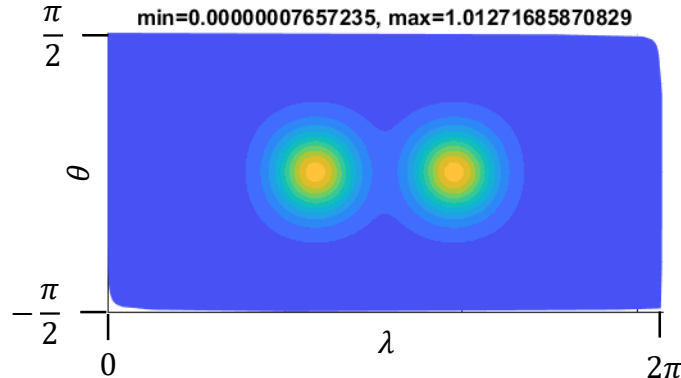


min=0.00000870799572, max=0.98992711564059

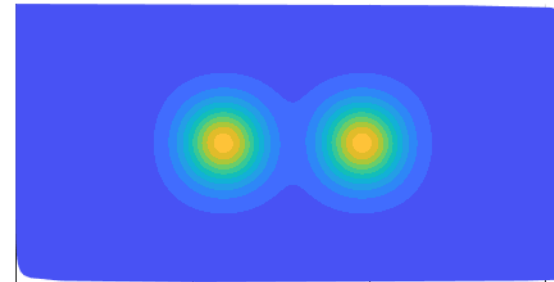


$N = 192^2$
 $\approx 1.1^\circ$

min=0.00000007657235, max=1.01271685870829



min=0.00000012431515, max=0.99616671662682



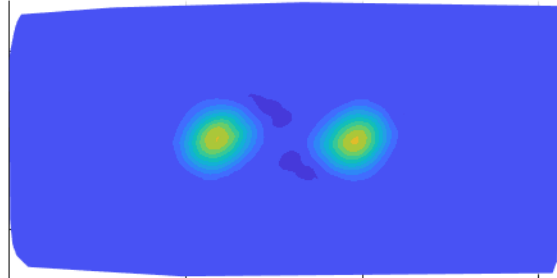
CB, Velocity Case 2

Unlimited

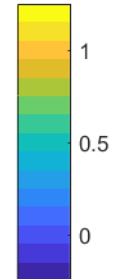
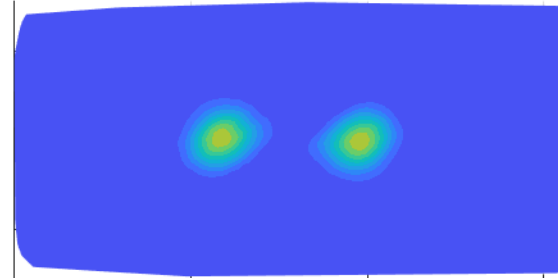
Limited

$N = 48^2$
 $\approx 4.4^\circ$

min=-0.08003766337462, max=0.87309378889769

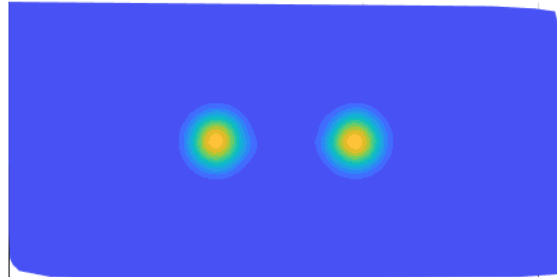


min=0.00000000000000, max=0.80940960356351

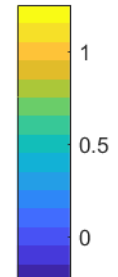
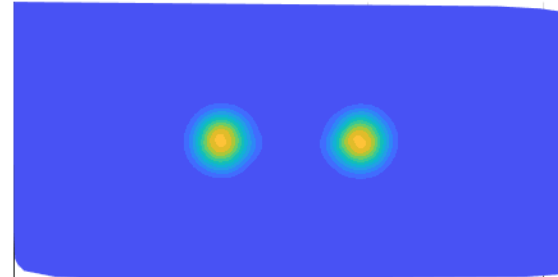


$N = 96^2$
 $\approx 2.2^\circ$

min=-0.04464567496036, max=1.02497099158292

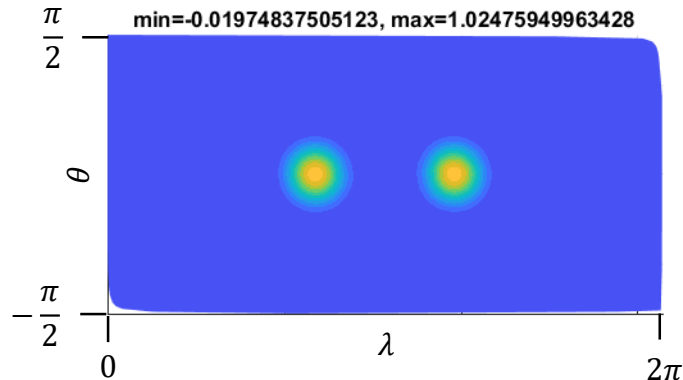


min=0.00000000000000, max=0.97431850267832

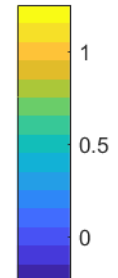
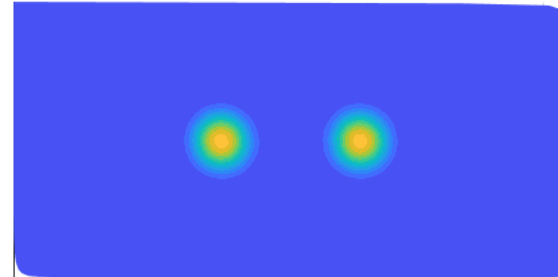


$N = 192^2$
 $\approx 1.1^\circ$

min=-0.01974837505123, max=1.02475949963428



min=0.00000000000000, max=0.99118591213709



SC, Velocity Case 2

Unlimited

Limited

$N = 48^2$
 $\approx 4.4^\circ$

min=-0.17961755562610, max=1.18859422521987

min=0.00000000000000, max=0.97455669006563

$N = 96^2$
 $\approx 2.2^\circ$

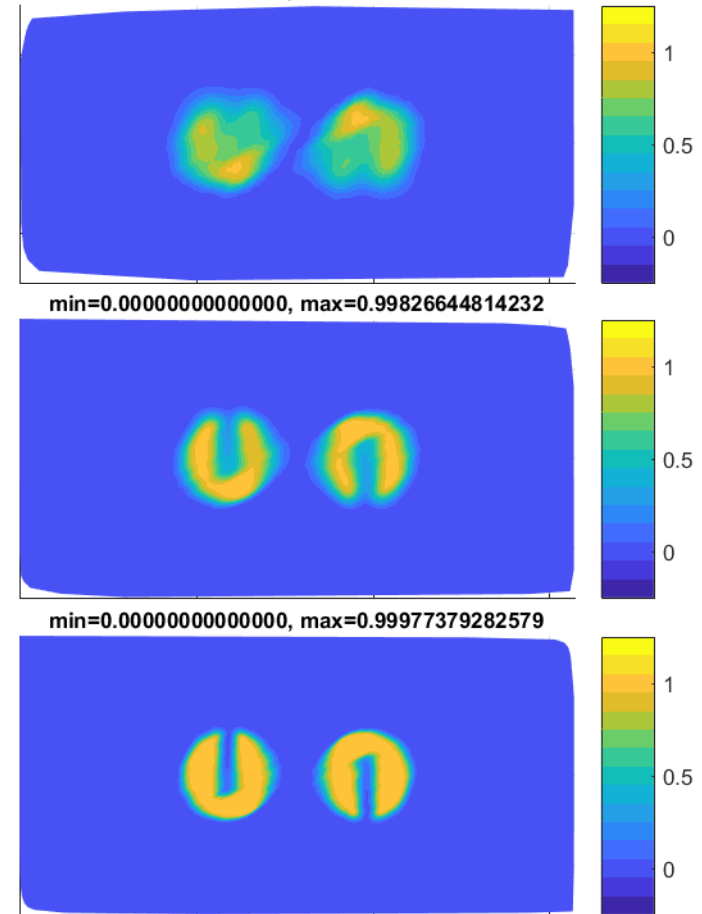
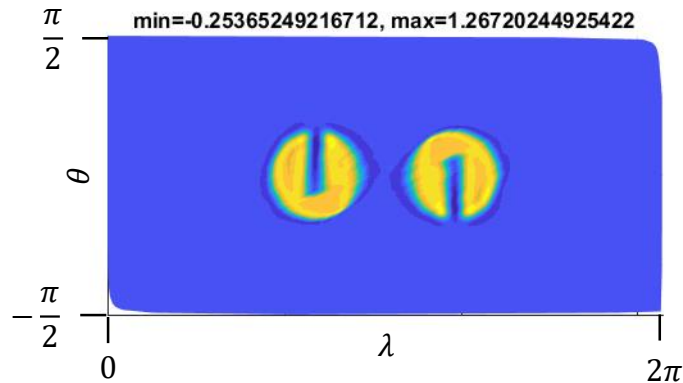
min=-0.16301110615044, max=1.26414651431357

min=0.00000000000000, max=0.99826644814232

$N = 192^2$
 $\approx 1.1^\circ$

min=-0.25365249216712, max=1.26720244925422

min=0.00000000000000, max=0.99977379282579



Advantages of Semi-Lagrangian Transport

- Large Time steps $\left(\frac{\max\|\mathbf{u}\|\Delta t}{\Delta x} \gg 1\right)$
- Simple governing equation and solution algorithm
 - $\frac{Dq}{Dt} = 0 \quad \left(\frac{D}{Dt} = \frac{\partial}{\partial t} + \mathbf{u} \cdot \nabla\right)$
 - q is constant along flow trajectories
 - No spatial derivatives
 - (1) time-step for departure points, (2) interpolate for new values of q
- No need for hyperviscosity
 - High frequencies automatically damped by repeated interpolation
 - Numerical solutions remain bounded even if the node set is poorly distributed
- Simple limiter to reduce oscillations and preserve bounds

Shallow Water Model

Governing Equations:

$$\frac{\partial \mathbf{u}}{\partial t} = -\mathbf{u} \cdot \nabla \mathbf{u} - f(\hat{\mathbf{k}} \times \mathbf{u}) - g \nabla h,$$

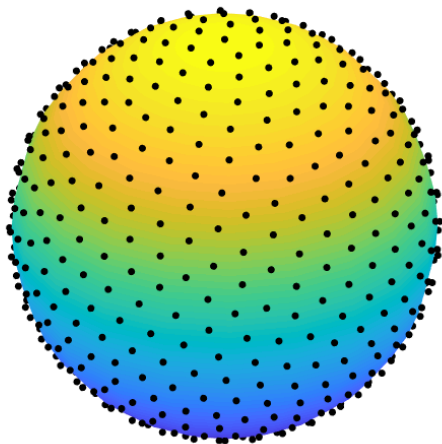
$$\frac{\partial h^*}{\partial t} = -\mathbf{u} \cdot \nabla h^* - h^*(\nabla \cdot \mathbf{u}).$$

- $f = 2\Omega \sin \theta$, where Ω = angular velocity of Earth, θ = latitude
- $\hat{\mathbf{k}}$ is the unit normal to the sphere
- g is gravitational acceleration
- $h = h_s(x, y, z) + h^*(x, y, z, t)$ is the depth of the fluid

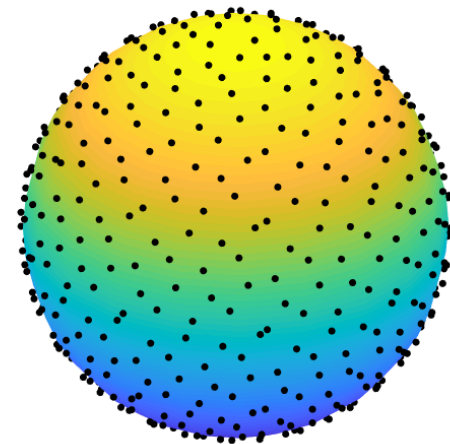
Note: The velocity \mathbf{u} is adjusted after every Runge-Kutta stage to remain tangent to the sphere ($\mathbf{u} \leftarrow \mathbf{u} - (\mathbf{u} \cdot \hat{\mathbf{k}})\hat{\mathbf{k}}$)

Nodes

Maximum Determinant (MD)



Hammersley



Shallow Water Test Cases

- Taken from Williamson et al, JCP 1992
 - Steady-state smooth flow
 - $h_s = 0$
 - Exact solution known
 - Flow over an isolated mountain
 - h_s is a cone-shaped mountain centered at $(\lambda, \theta) = \left(-\frac{\pi}{2}, \frac{\pi}{6}\right)$
 - Exact solution unavailable
 - Rossby-Haurwitz Wave
 - \mathbf{u}_0 and h_0 satisfy the barotropic vorticity equations
 - $h_s = 0$
 - Exact solution unavailable

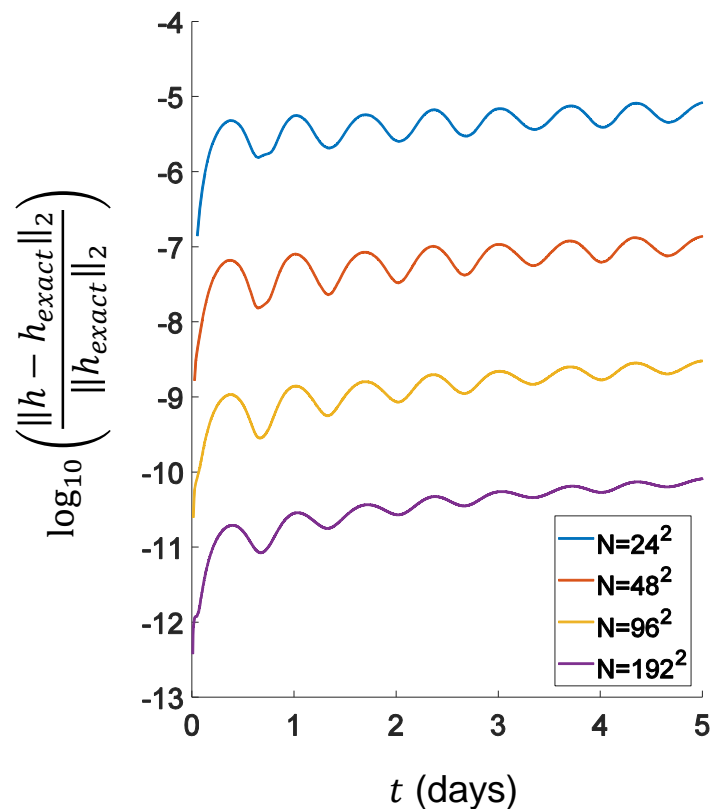
Parameters for Shallow Water Tests

- Derivative approximations $\left(\frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z}\right)$
 - $\phi(\underline{x}) = \|\underline{x}\|^2 \log \|\underline{x}\|$
 - Polynomials up to degree 5
 - Stencil size 42 (twice as many RBFs as polynomials)
- Hyperviscosity (Δ^3)
 - $\phi(\underline{x}) = \|\underline{x}\|^7$
 - Polynomials up to degree 5
 - Stencil size 42
 - Parameter $\gamma = 2^{-12} \approx 2.4 \times 10^{-4}$
- Time Stepping (3 stage, 3rd order Runge-Kutta)

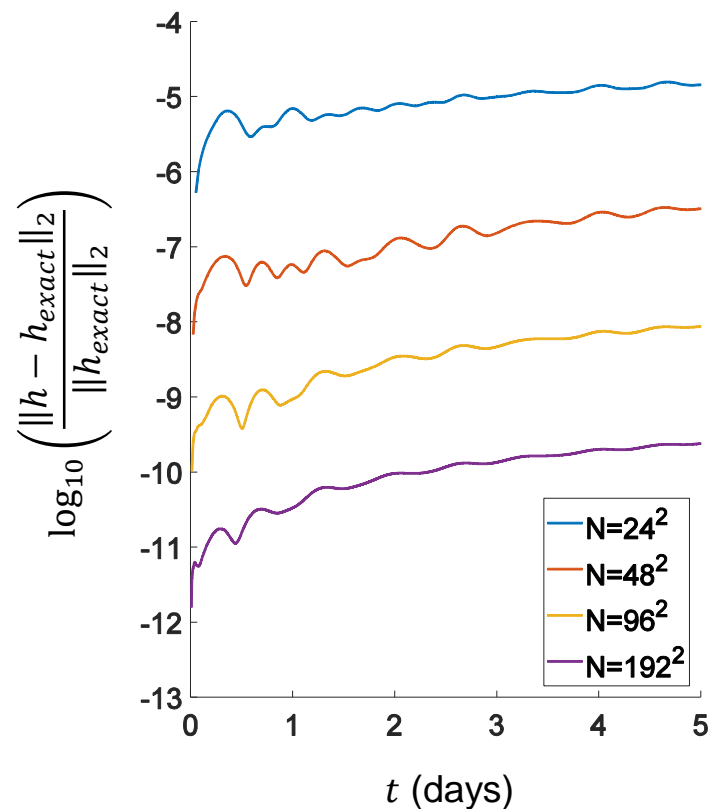
	$N = 24^2$	$N = 48^2$	$N = 96^2$	$N = 192^2$
Δt (minutes)	36	18	9	4.5

Error Growth in Time

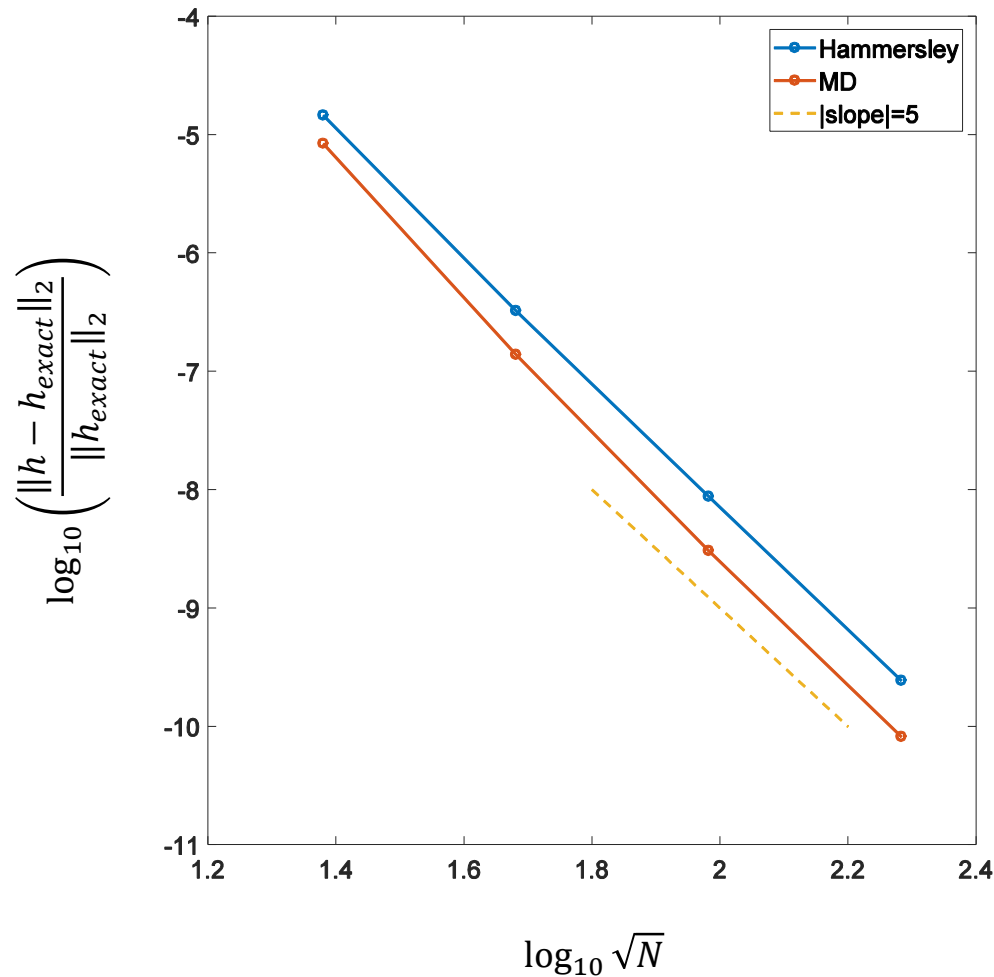
MD



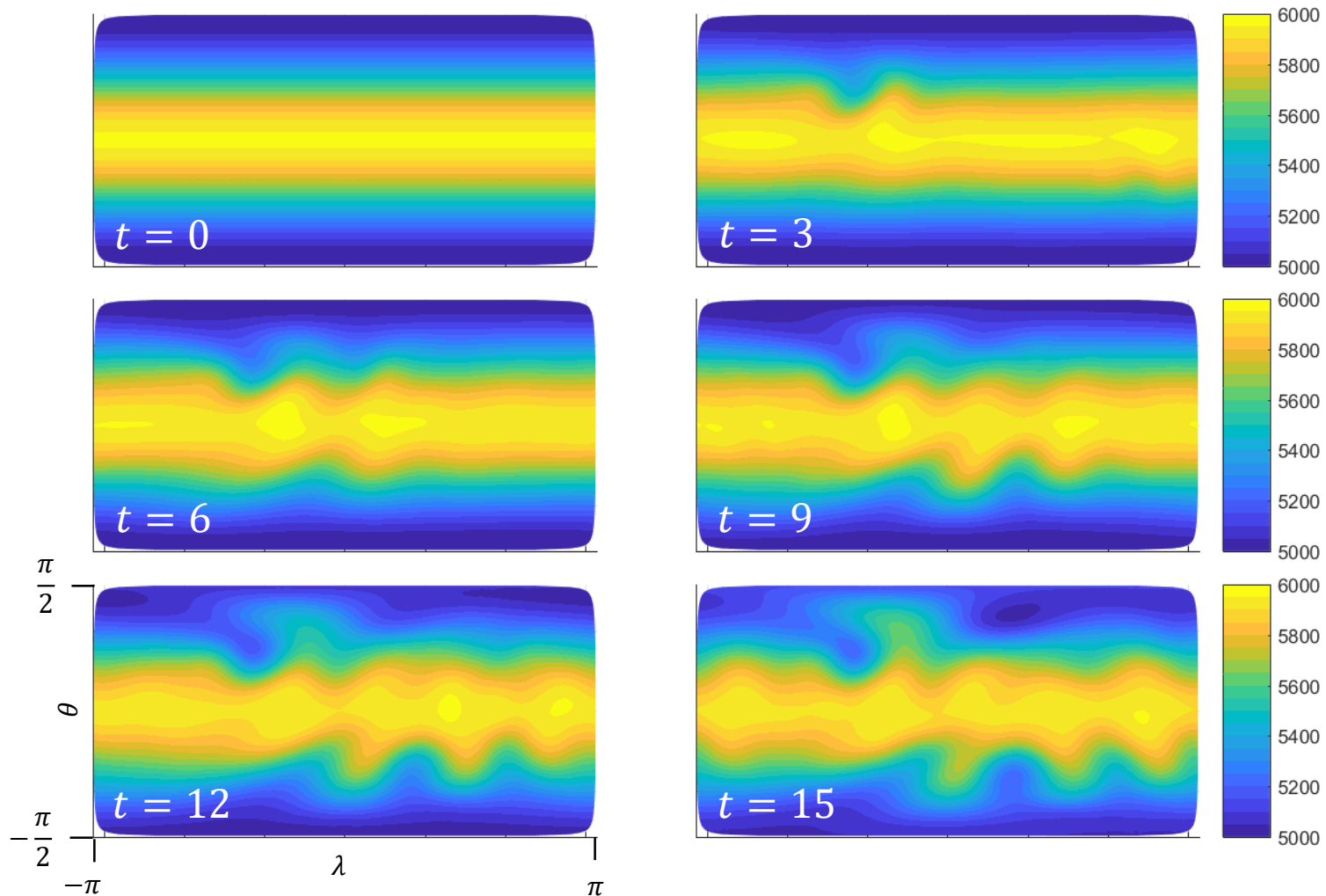
Hammersley



Convergence Verification ($t = 5$)



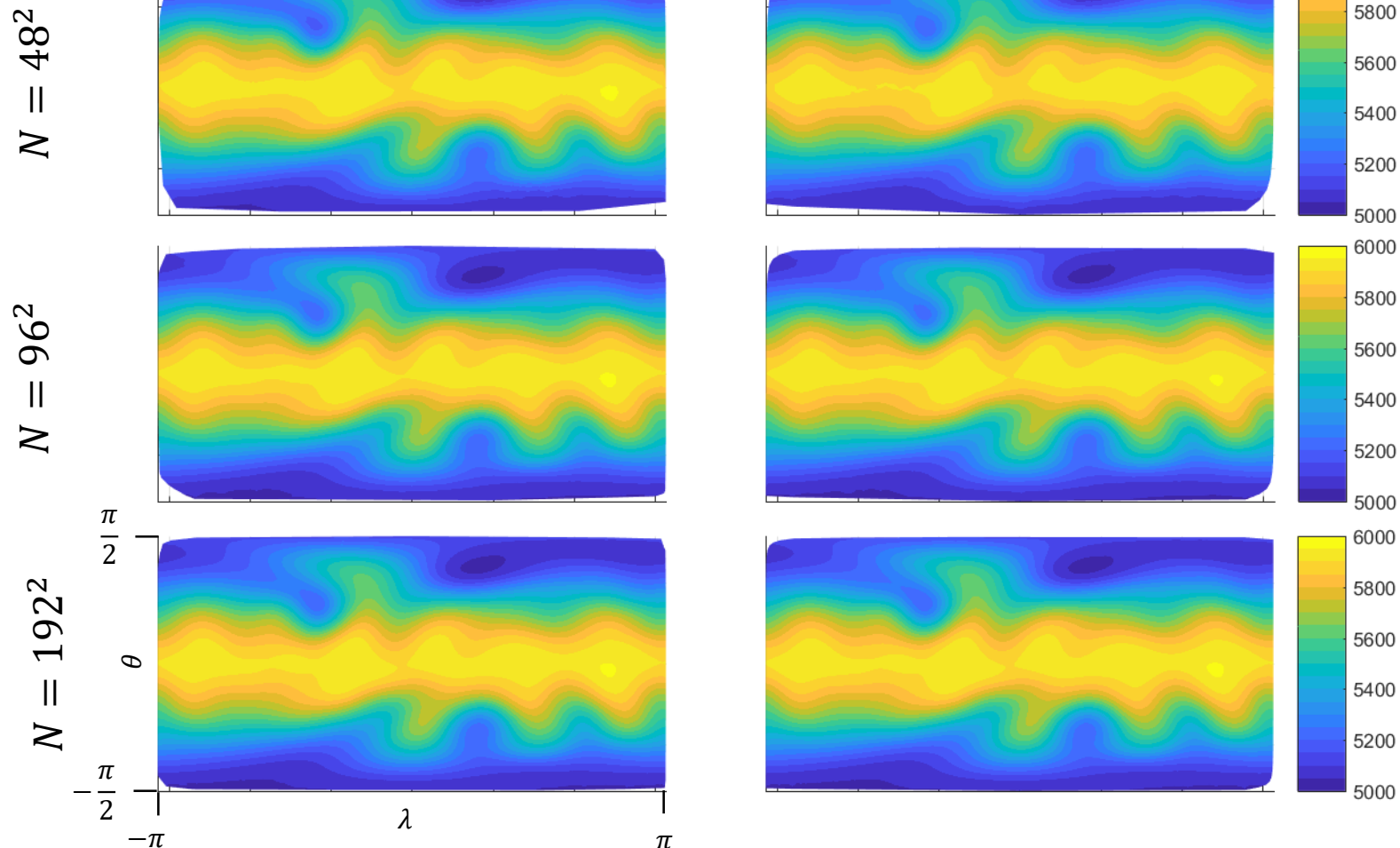
Time Snapshots, Isolated Mountain



Flow over Mountain, $t = 15$

MD

Hammersley

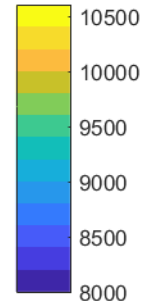
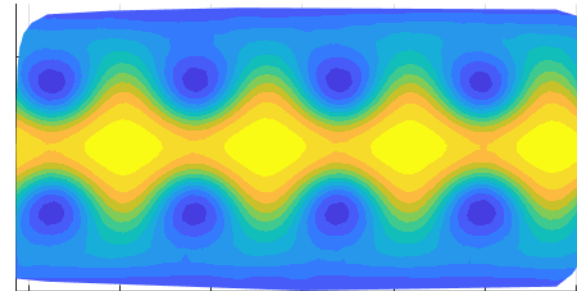
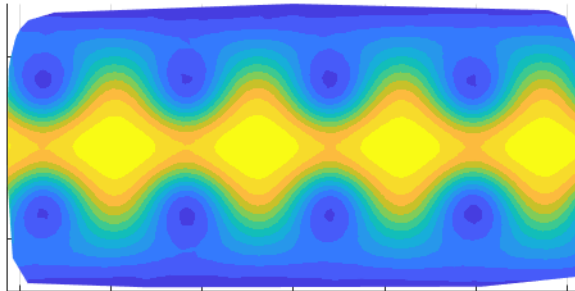


Rossby-Haurwitz, $t = 14$

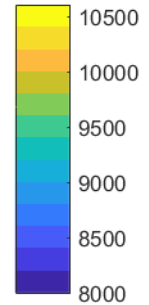
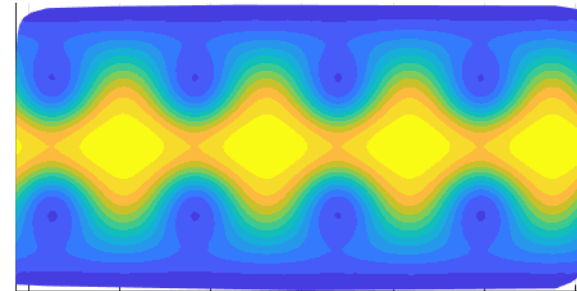
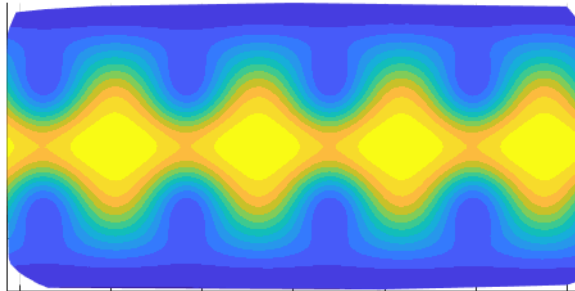
MD

Hammersley

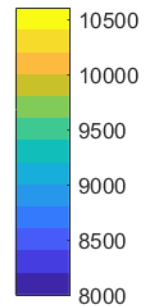
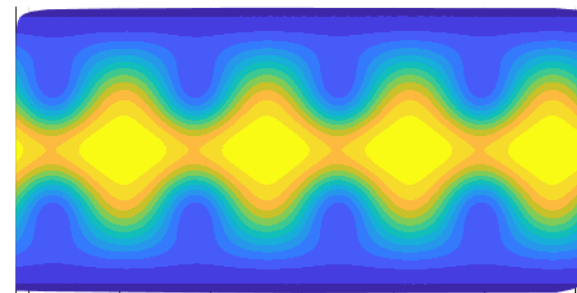
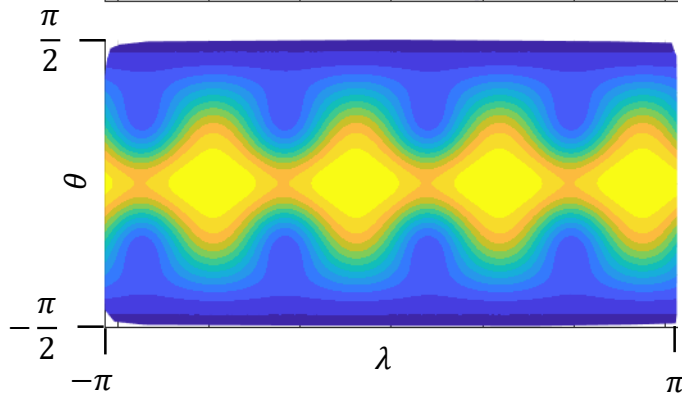
$N = 48^2$



$N = 96^2$



$N = 192^2$



Strengths of PHS RBF-FD

- Simple and accurate on the sphere
- Local and well suited for parallel computations
- Free from coordinate singularities
 - Discretize directly from Cartesian equations
- Geometrically flexible
 - Does not require a mesh
 - Static Node Refinement
 - Dynamic Node Refinement
- Robust
 - Same configuration (basis, stencil-size, hyperviscosity parameter) runs on a wide variety of node-sets and test problems
 - $\|x\|^2 \log\|x\| + p5 + n42$ for first derivative approximations

Future Work

- Transport
 - 3D test cases on spherical shell (DCMIP test cases)
 - More sophisticated fixer/limiter procedure
 - Reduce parallel communication
- Shallow water equations
 - Quantitative comparison to other methods
 - Additional tests on the sphere from Williamson et al, JCP 1992
 - Forced nonlinear system with a translating Low
 - Evolution of highly nonlinear wave
- Nonhydrostatic Dynamical Core for climate/weather
 - 2D benchmarks in Cartesian geometry with topography
 - Fully 3D without using a terrain-following coordinate transformation
 - Eulerian dynamics, semi-Lagrangian transport with fixer/limiter

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